

# Building a Telescope 2 km Underground

Jason Detwiler

James Loach

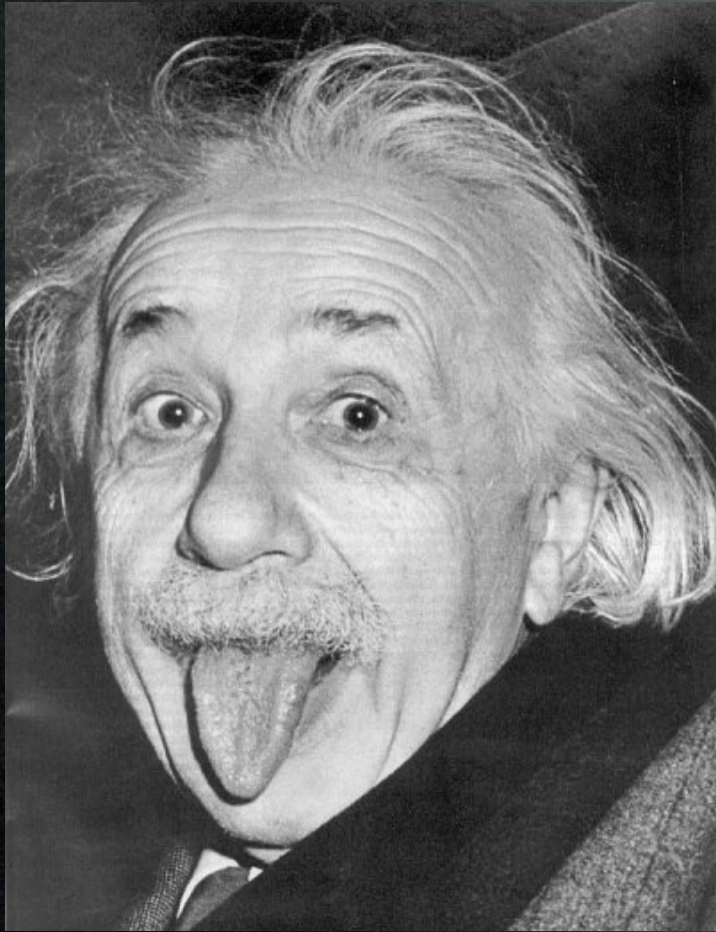
Alan Poon

Gersende Prior

Lawrence Berkeley National Lab



# Physics



- We are physicists, we study the properties of energy, matter, space and time.
- **Albert Einstein** was a famous theoretical physicist. Theoretical physicists try to explain physical phenomena with mathematical models.

$$E = mc^2$$



# Physicists?

- But we are **experimental physicists**. We build experiments and try to test the predictions made by the theoretical models.
- **Ernest O. Lawrence** was a famous experimental physicist. Lawrence Berkeley National Lab and Lawrence Hall of Science were named after him.





# “Light”

- What is the similarity between optical light (which your eyes are sensitive to), radio-wave (FM radio), and X-ray?
- They are all “particles of light” (photons). The only difference between them is their energy. Our eyes are sensitive to photons of relatively low in energy.



# Telescopes



- Instruments that allow you to observe distant stellar objects
- The most familiar ones are optical telescopes.
- They are usually ground-based, and are usually located at mountain tops.





# Telescopes

- There are telescopes that observe photons in different energy regimes

Space-based



Chandra (X-ray)

Ground-based



Very Large Array, NM  
(Radio)

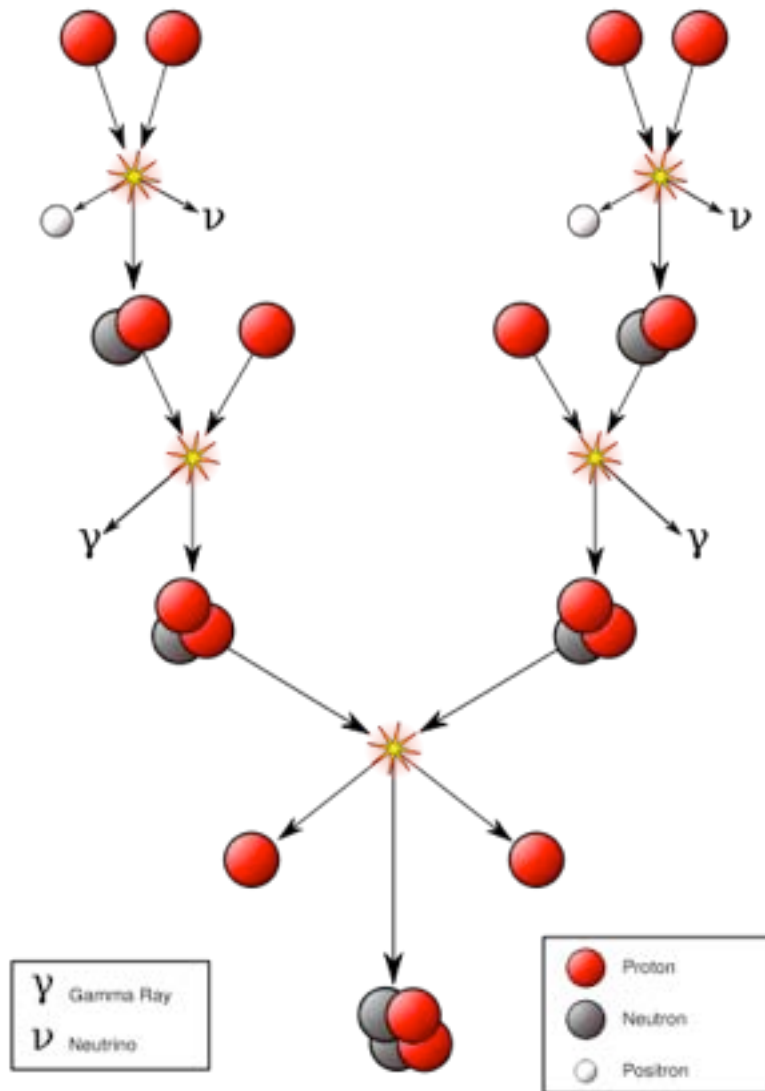


# Telescopes

- But telescopes do not necessarily detect photons.
- Photons are not good probes to the inner workings of stellar objects.
- Why not?



# Our Sun



- Our Sun is a **star**. Stars generate energy by **nuclear fusion**, in which protons are fused to form Helium nuclei (2 protons and 2 neutrons) through a sequence of nuclear reactions in the core.
- Photons from these processes take tens of thousands of years to travel from the solar core to the surface.
- Once the photons have reached the solar surface, they take a further 8 minutes to reach us on Earth.



# Our Sun

- So the sunlight we see is at least 10000 years old!
- So how do we know the Sun generates its energy through nuclear fusion in its core?
- Neutrinos!



# Neutrinos

- Neutrinos are byproducts of the nuclear fusion process.
- Unlike photons, neutrinos do not interact with matter very much. Neutrinos take almost no time to travel from the solar core to the surface.
- Like photons, neutrinos are fundamental particles in nature.



# Neutrinos

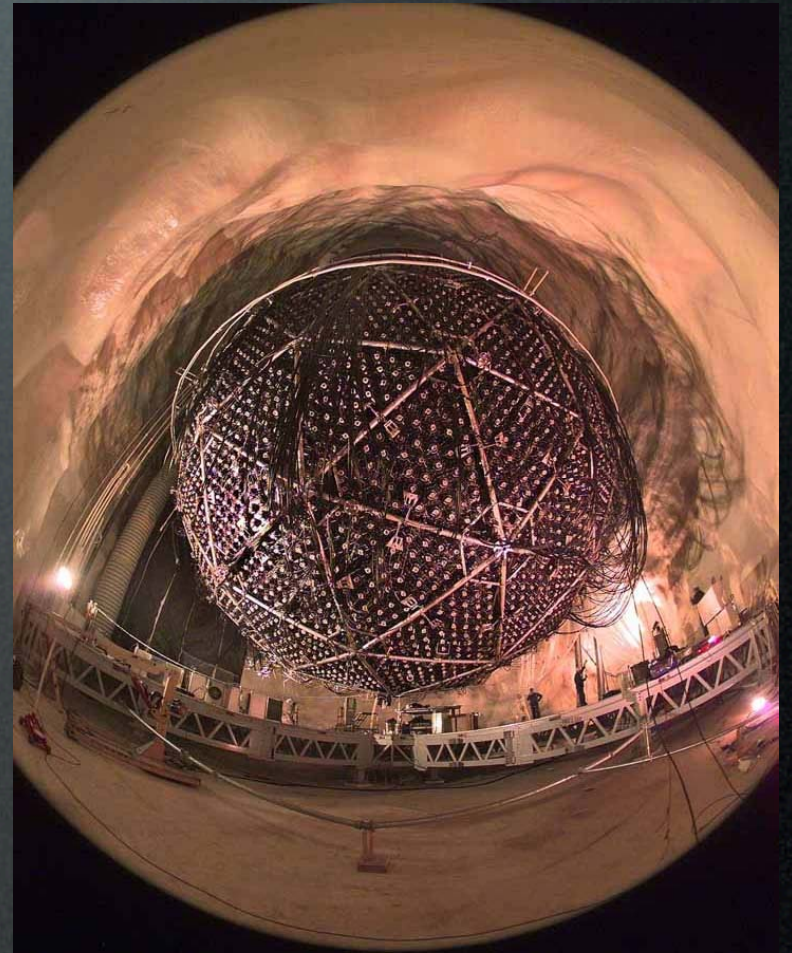
- Neutrinos are byproducts of the nuclear fusion process.
- Unlike photons, neutrinos do not interact with matter very much. Neutrinos take almost no time to travel from the solar core to the surface.
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How do you detect neutrinos?



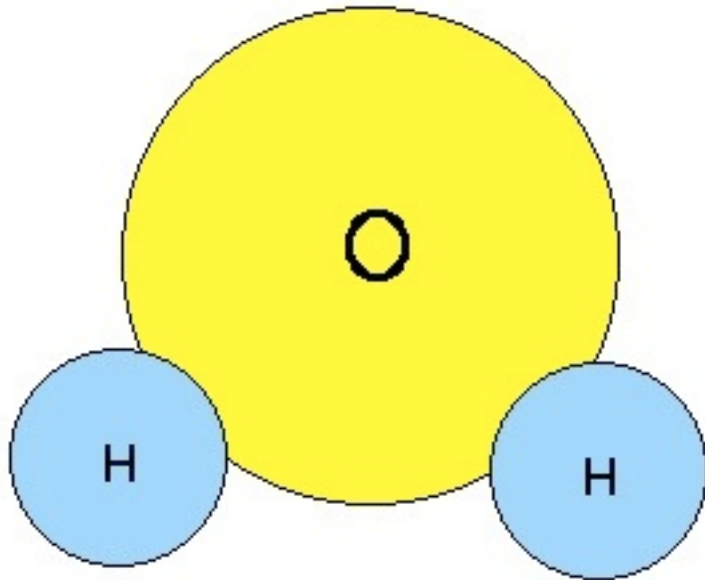
# Sudbury Neutrino Observatory (SNO)

- SNO is a neutrino telescope
- It is located in a Canadian mine more than a mile underground.
- Neutrinos interact with the **heavy water** inside SNO, and create flashes of light.
- This telescope has about 10000 light sensors to detect these flashes.



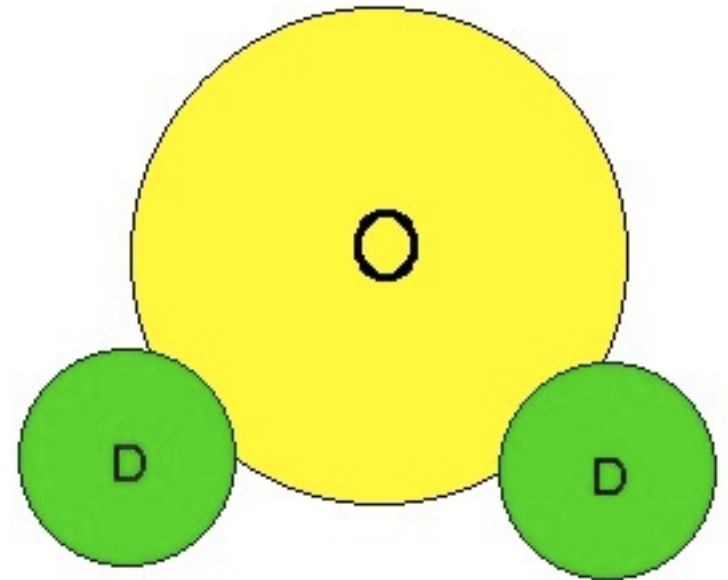


# Water vs Heavy Water



Water

**H**: hydrogen (1 proton)



Heavy Water

**D**: deuterium  
(1 proton + 1 neutron)

Deuterium is good at stopping neutrinos

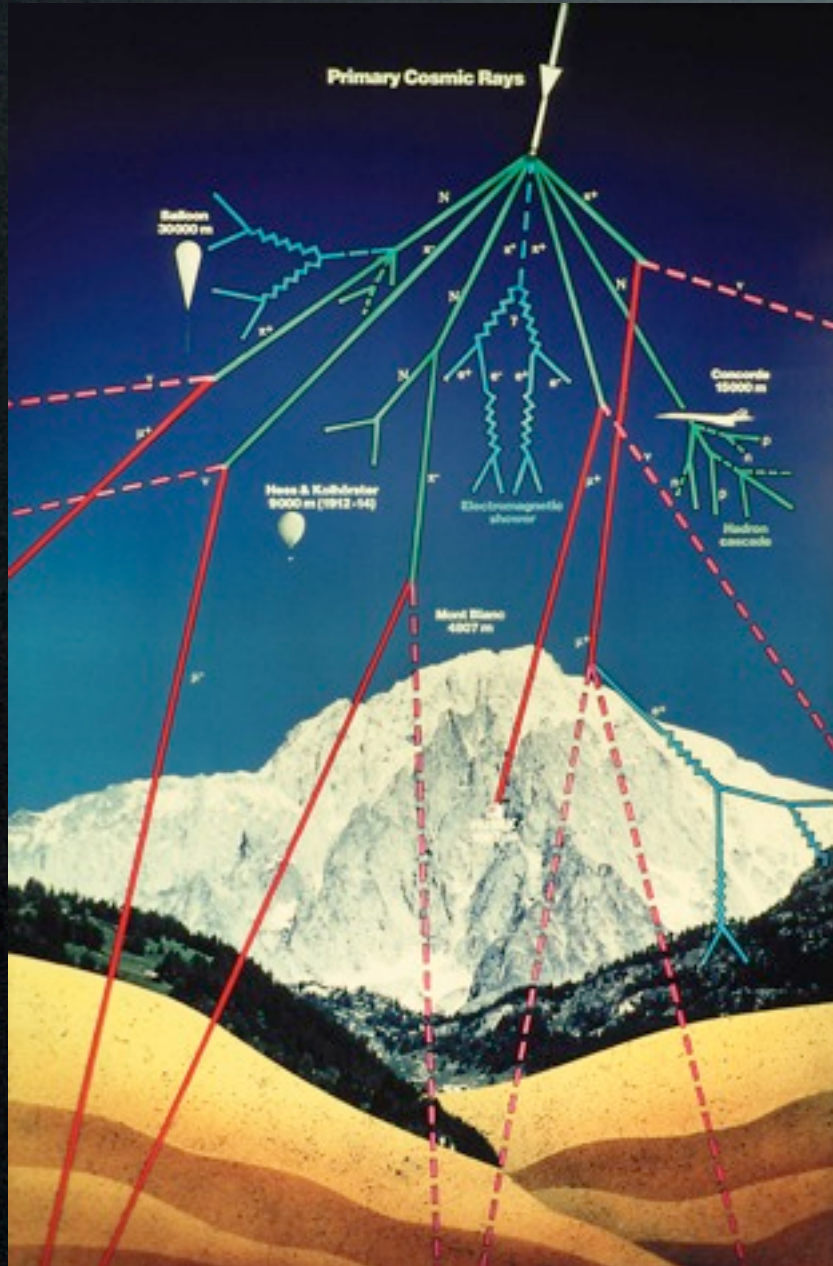


# Why did you build a telescope underground?

- Neutrinos can go through almost everything. SNO catches only a few neutrinos from the Sun every day.
- But there are **cosmic rays** from space. The Earth can stop cosmic rays, but not neutrinos. We can find the neutrinos more easily if the telescope is deep underground.



# What are cosmic rays?



- High energy particles (mostly protons) from our solar system or other galactic and extra-galactic sources bombard our Earth's upper atmosphere.
- They interact with nuclei in the atmosphere (nitrogen, oxygen...) and create a shower of other particles.
- A significant fraction of these secondary particles are **muons**, a heavier cousin of electrons.
- You will make measurements of these muons later.



# Can you find Alan?



No, there are too many people!



# Can you find Alan?



Maybe, if you look very carefully.



# Can you find Alan?



Yes, because there are fewer people in the picture.  
If there are fewer cosmic rays, it is easier to find  
neutrinos too.



# SNO

- SNO is located in Sudbury, Canada.
- It is a joint project between Canada, US and the UK.
- The heavy water (worth about \$300 Million) was on loan from the Canadian government.





# A SNO Tour



This is the entrance to the mine.



# A SNO Tour



We take the **elevator** (aka cage) down to the mine.



# A SNO Tour



After we get off the elevator, we walk a mile to our laboratory.



# A SNO Tour



The blue door is the entrance to the laboratory.



# A SNO Tour



Then we take off our miner clothing and shower. We have to be very **clean** when we work on SNO. The lab is a cleanroom



# A SNO Tour



This is the lunch room where we take a break after the shower!



# A SNO Tour



This is the water system that circulates and purifies the heavy water



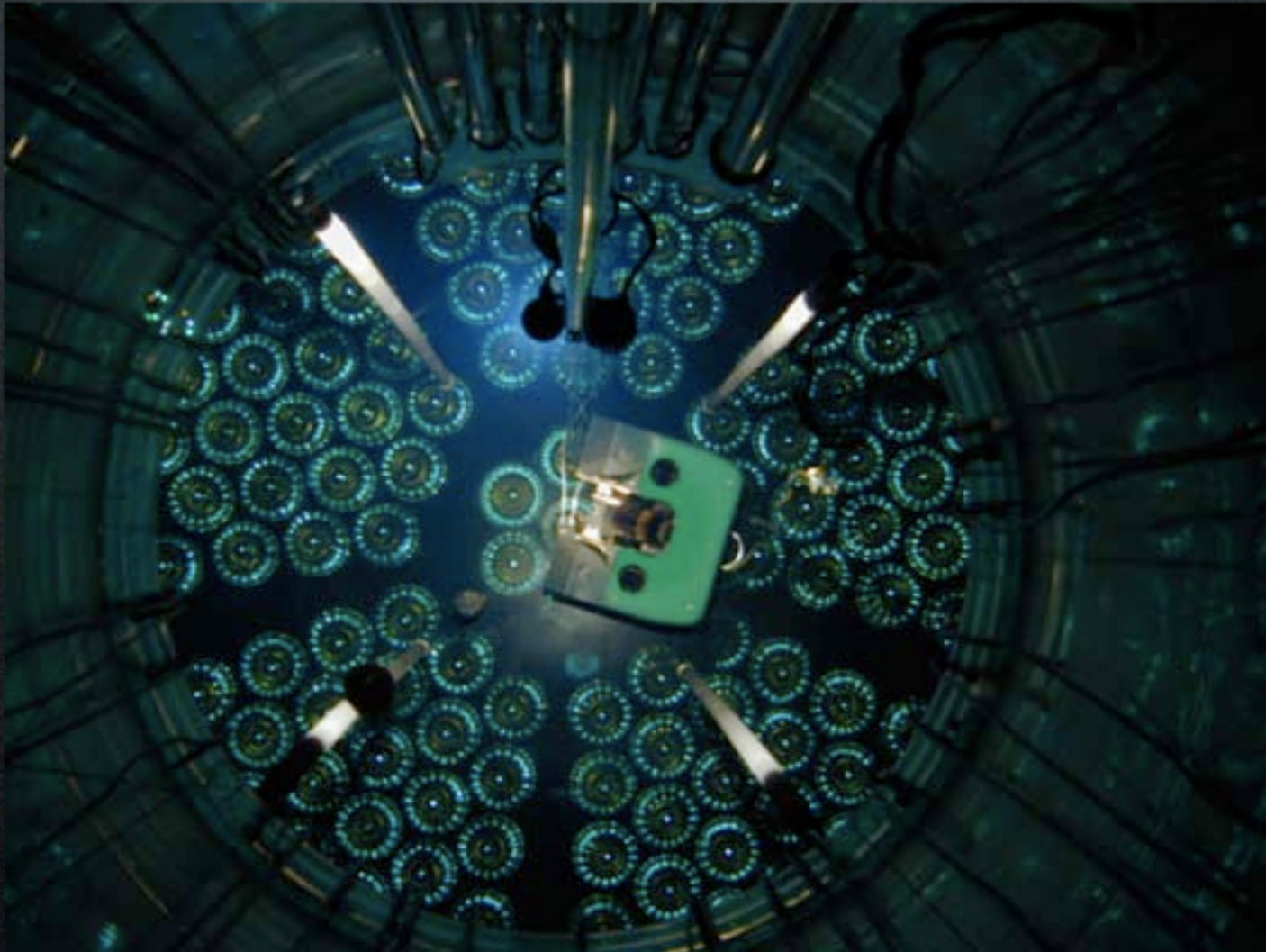
# A SNO Tour



We have to check on SNO sometimes.  
We use a rubber boat.



# A SNO Tour



We sent a submarine into SNO to install certain components.



# A SNO Tour



Young students pilot the submarine because they are good at video games.



# A SNO Tour



At the end of the day, we walk back to the elevator.  
It takes us back to surface.



# What did we found?

- SNO observed neutrinos from the Sun.
- SNO results show that computer models of the Sun predict the correct total number of neutrinos from the fusion process.
- Theoretical physicists thought neutrinos had zero mass (like photons). SNO proved them wrong.



# SNO

- Show event display here



# Muon Measurement

- To determine how the muon flux (# of muons per unit area) depends on the zenith angle
- Adjust H
- Count for 3 min. at each H
- Plot the results
- How would you explain the results?

